A 3D rendering of the VIPER rover on the lunar surface. The rover is a six-wheeled rover with a large solar panel on its left side and a mast with a camera and light at the top. The background shows a dark lunar landscape with a bright light source on the horizon.

VIPER

Overview

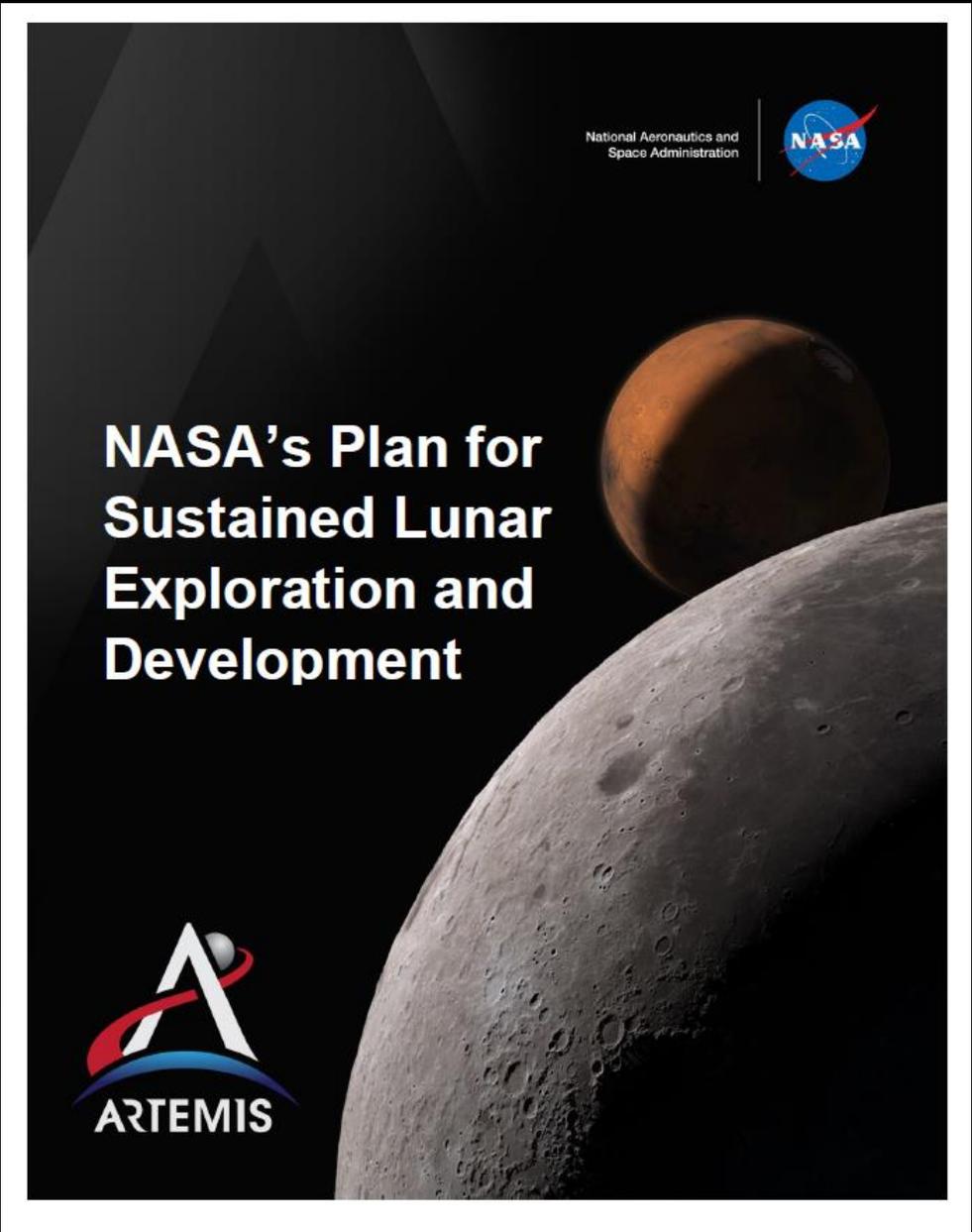
Volatiles Investigating Polar Exploration Rover

Daniel Andrews, PM
European Lunar Symposium
MAY 12-14, 2020

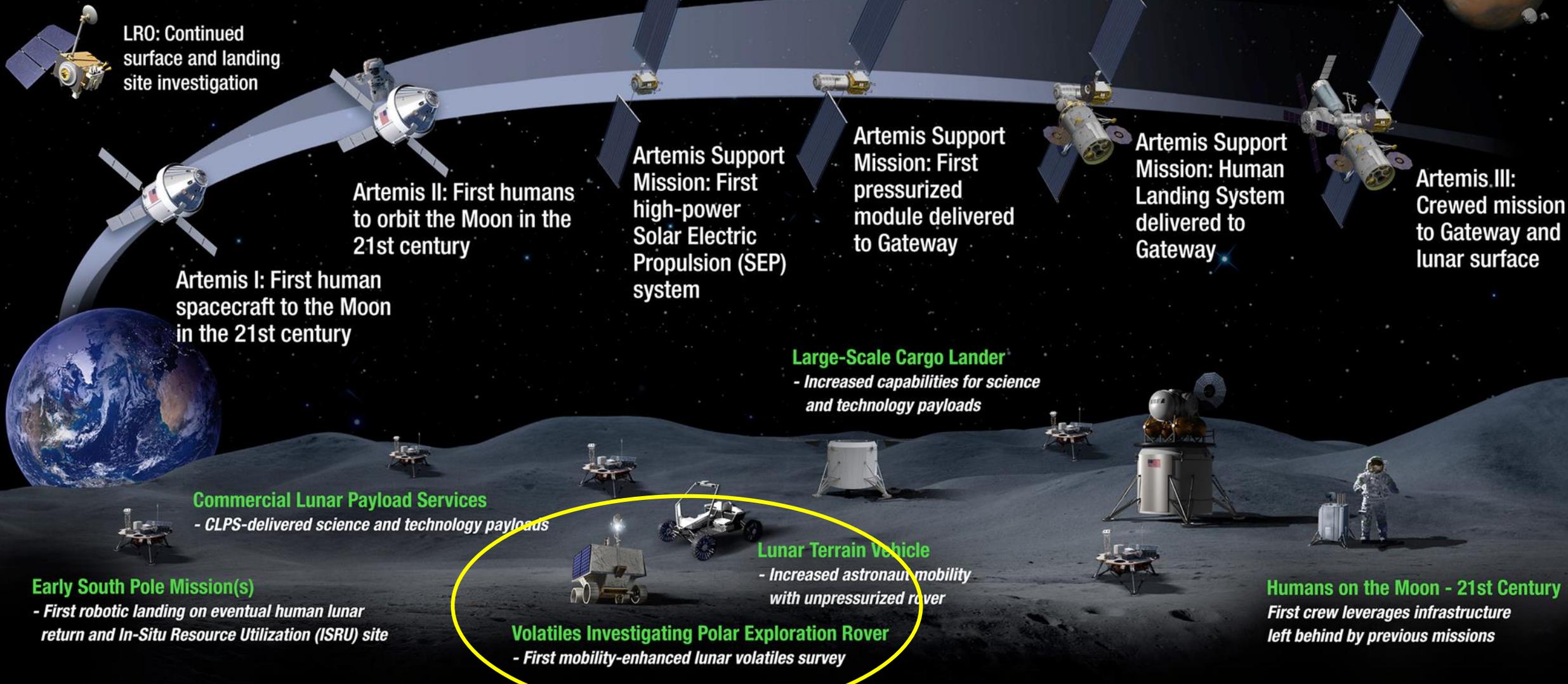
VIPER & Artemis

Our return to the Moon begins with robots. NASA's new Commercial Lunar Payload Services (CLPS) initiative will deliver its next robotic lunar rover, the Volatiles Investigating Polar Exploration Rover (VIPER)

VIPER will conduct science investigations of the lunar volatiles at the Moon's South Pole. The data produced by VIPER will inform future in-situ resource utilization (ISRU) technologies.



Humans Return by 2024



LUNAR SOUTH POLE TARGET SITE

2020

2024



Why VIPER?

Direct measurement of polar volatiles

- *LCROSS ground-truthed the water* - VIPER will reveal the **lateral / vertical distribution** and **physical state / composition** of the volatiles
- *Q: Are some polar regions better than others (feasibility, economics, safety)?*

Enables research into In-Situ Lunar Resources

- VIPER will **Build Lunar resource maps**, steering the future commercial marketplace
- Understand **ore grade availability of lunar volatiles** for human sustainment and fuel

Where will VIPER explore?

VIPER will explore four “**Ice Stability Regions**” (ISRs)*:

- **Surface**: Ice expected stable at the surface - Permanently Shadowed Regions, (PSRs)
- **Shallow**: Ice expected stable within 50cm of surface
- **Deep**: Ice expected stable between 50-100 cm of the surface
- **Dry**: Ice *not* expected (top meter to be *too warm* to be stable)

**ISRs based on the predicted thermal stability of ice with depth*

VIPER Performance Specs



- **Mass:** ~475kg (1050lbs) **Power (peak):** ~450W
- **Comms (DTE¹):** X-band
 - 256kbps (high-gain) Moon-to-Earth (as high as 4Mbps)
 - 2kbps (omni) Earth-to-Moon
 - Ground: DSN 34m dishes: Canberra, Goldstone, Madrid
- **Dimensions:** 1.5m x 1.5m x 2.5m (5ft x 5ft x 8ft)
- **Top Speed:** 20cm/s (0.5MPH)
- **Expected Cold Environment:** -230degC (-382degF)
- **Prospecting Speed:** 10cm/s (0.25MPH)
- **Distance Travelled (goal):** 20km (~12mi)
- **Lunar delivery:** CLPS² commercial contract

¹ DTE = Direct-To-Earth

² CLPS = Commercial Lunar Payload Services



VIPER Science Specs

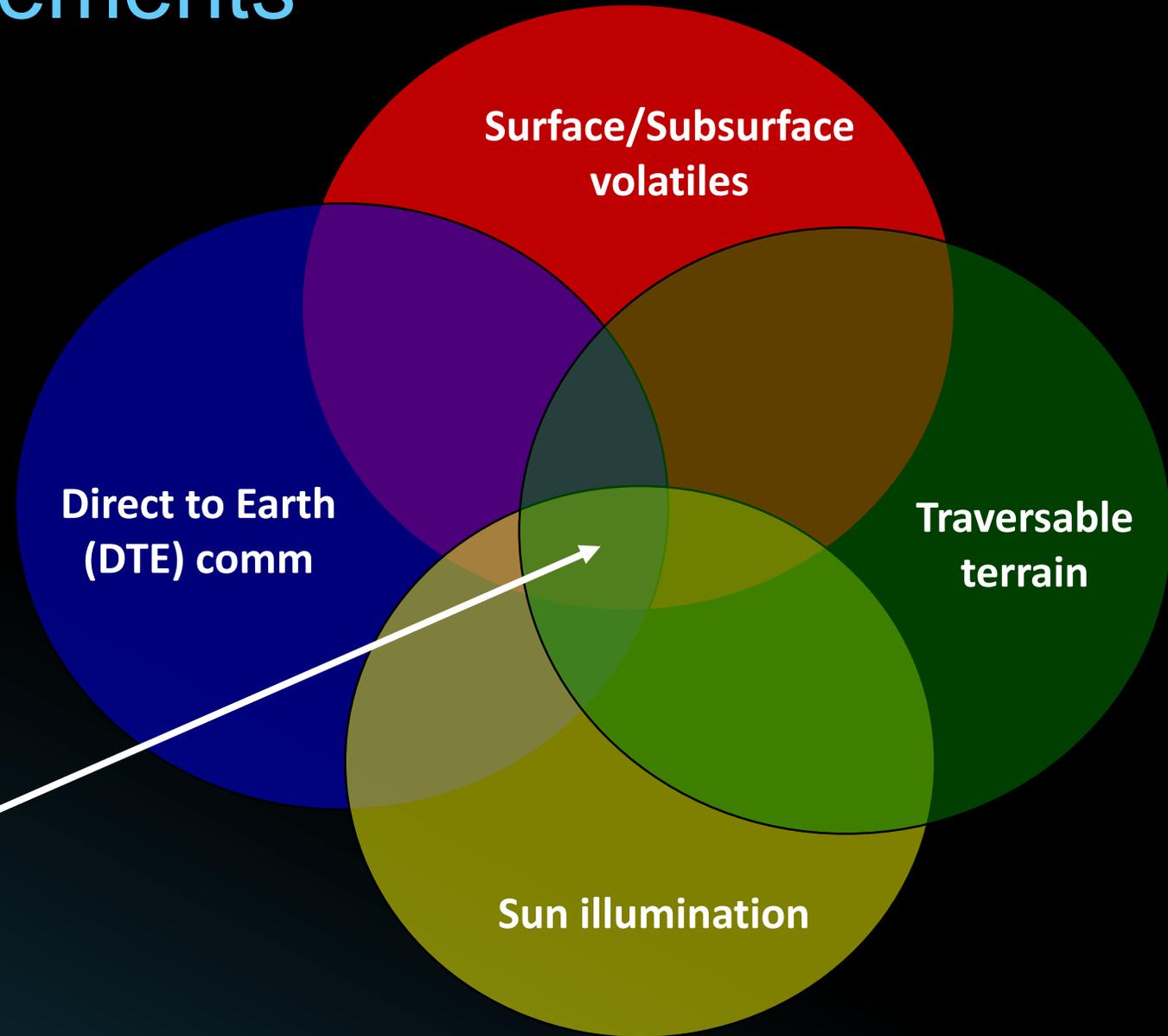
- **Mission Duration:** 100+ earth days
- **Instruments:** Neutron, Near-IR, and Mass Specs; 1m Drill
- **Detectable H₂O Concentration:** 0.5% (by weight)
- **Drill Depth:** 1m (~3ft)
- **# of Surface Assays (drill sites):** 18
- **Dark Survivability:** 96hrs (VIPER driving case)
- **PSR Working Duration:** 6hrs (Resource Prospector driving case)
- **Surface Traverse Plan baselined:** @CDR (Q2/FY21)

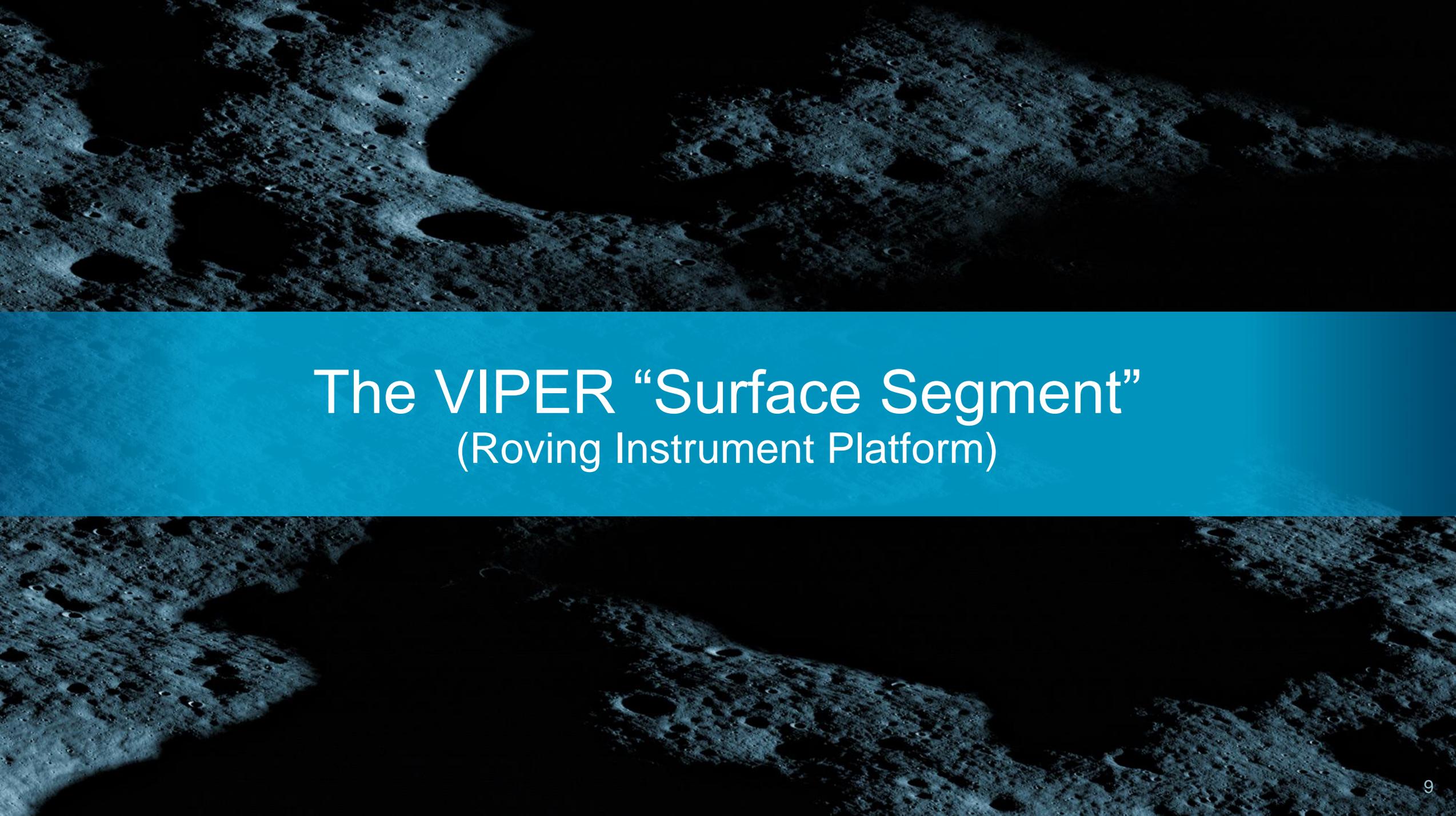
Landing Site Requirements

Good candidate polar landing sites meet these four criteria:

1. Surface/Subsurface Volatiles
2. Reasonable terrain for traverse
3. Direct view to Earth for communication
4. Sunlight for duration of mission for power

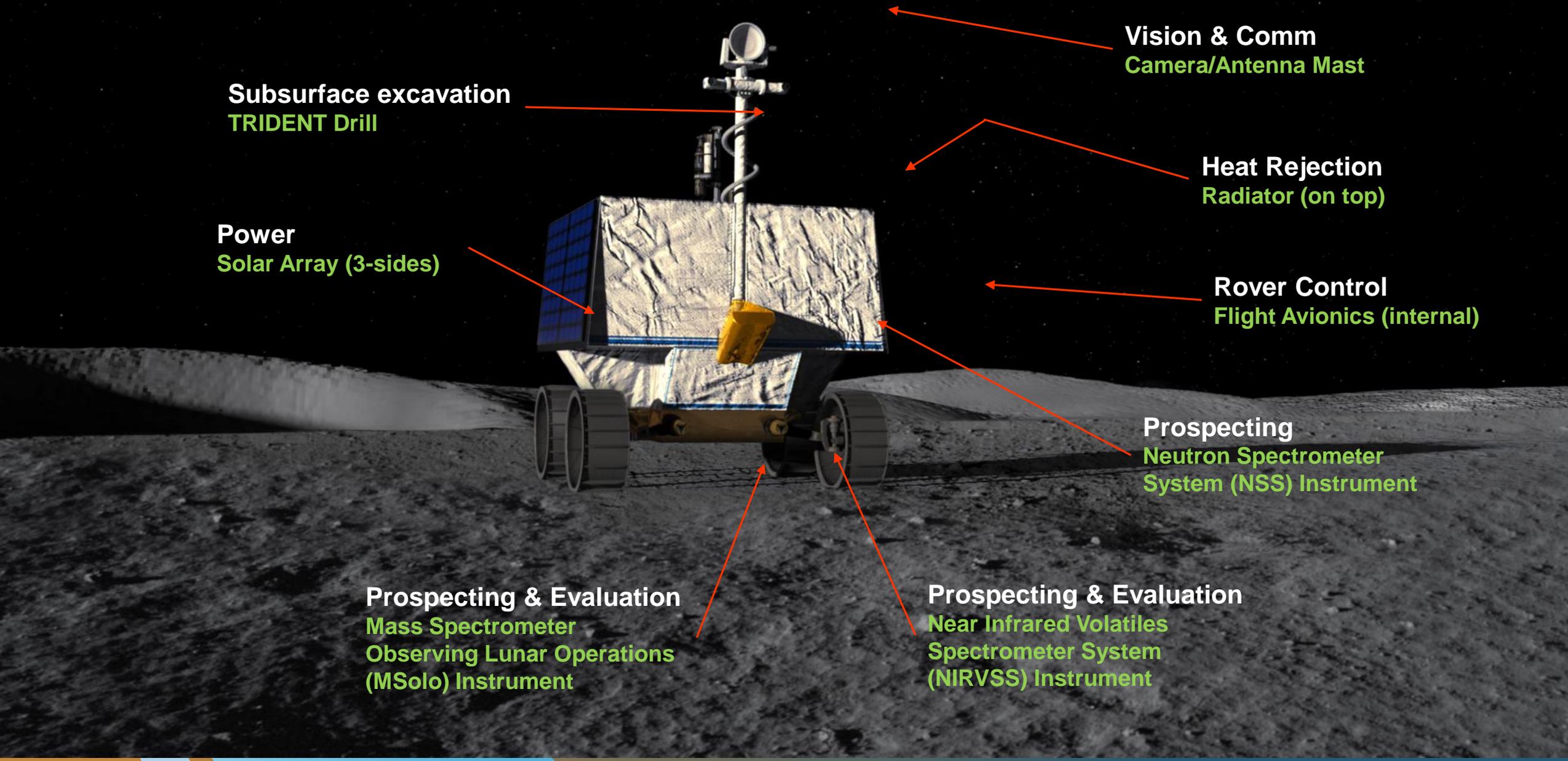
VIPER needs to find the intersection of these constraints



The background of the slide is a high-resolution image of the lunar surface, showing numerous craters of various sizes and depths. The surface is dark, with some highlights from the sun. A solid blue horizontal band runs across the middle of the image, containing the title text.

The VIPER “Surface Segment” (Roving Instrument Platform)

VIPER Surface Segment (Rover + Instruments)



Subsurface excavation
TRIDENT Drill

Power
Solar Array (3-sides)

Prospecting & Evaluation
Mass Spectrometer
Observing Lunar Operations
(MSolo) Instrument

Vision & Comm
Camera/Antenna Mast

Heat Rejection
Radiator (on top)

Rover Control
Flight Avionics (internal)

Prospecting
Neutron Spectrometer
System (NSS) Instrument

Prospecting & Evaluation
Near Infrared Volatiles
Spectrometer System
(NIRVSS) Instrument

Historical Planetary Rovers & VIPER

Driving on Other Worlds

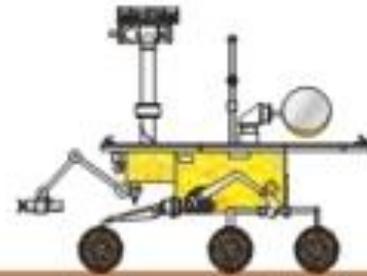
http://historicspacecraft.com/Probes_Mars.html

*includes instruments



Sojourner (1996):

- 0.6m x0.5m x0.3m
- 11kg
- Top Speed: 5cm/s



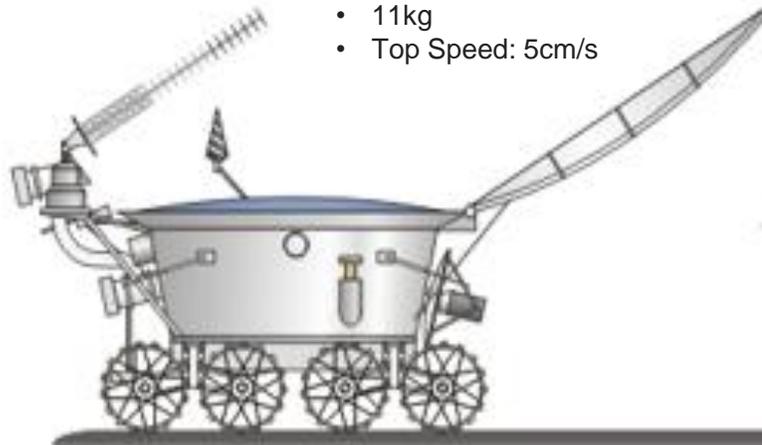
Spirit & Opportunity Mars Exploration Rover (2004):

- 1.6m x2.3m x1.5m
- 180kg*
- Top Speed: 5cm/s



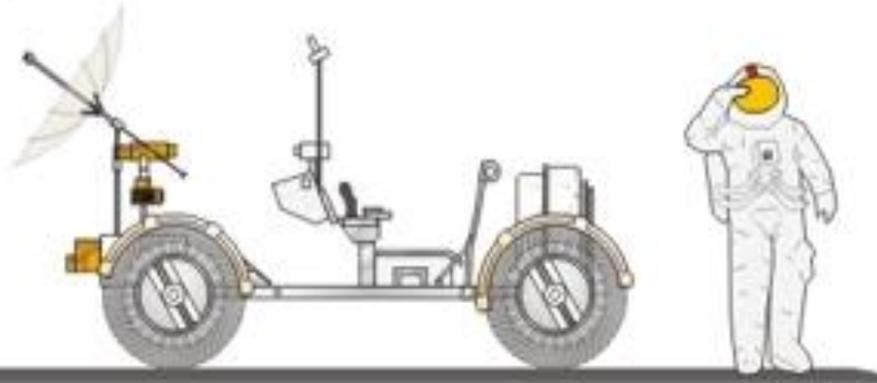
Curiosity Mars Science Laboratory (2011):

- 3.0m x2.8m x2.1m
- 900kg
- Top Speed: 4cm/s



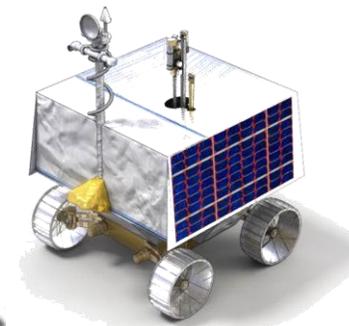
Lunokhod (1970/1973):

- 1.3M x1.6m x1.5m, 840kg
- Top Speed: 55cm/s



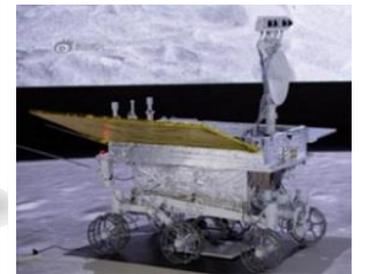
Lunar Roving Vehicle (1971/1972):

- 1.3M x1.6m x1.5m, 840kg
- Top Speed: 500cm/s



VIPER (2022):

- 1.5m x1.5m x2.0m, 300kg
- Top Speed: 20cm/s



Yutu (2013):

- 1.5m x1.1m, 140kg
- 5cm/s

VIPER Science Manifest

Neutron Spectrometer System (NSS) NASA-ARC

- **Prospects for hydrogen-rich materials while roving**, mapping the distributions
- Located on the front of the rover to have an unobstructed view of the lunar surface

Near InfraRed Volatiles Spectrometer System (NIRVSS) NASA-ARC

- **Prospects for surface water and oxide “frosts”**, as well as mineralogical context
- Located under the rover studying water/volatiles abundance while roving & drilling

Mass Spectrometer observing lunar operations (MSolo) NASA-KSC

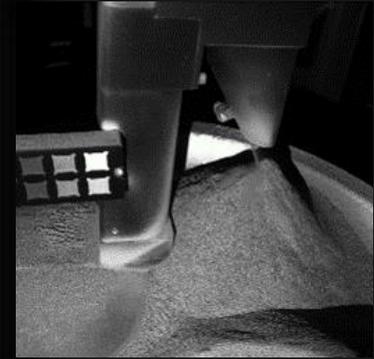
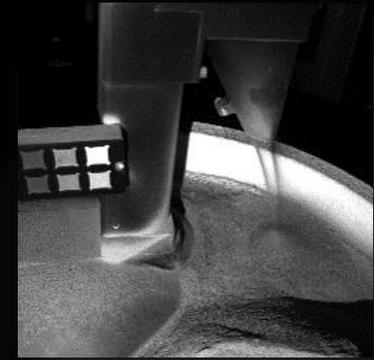
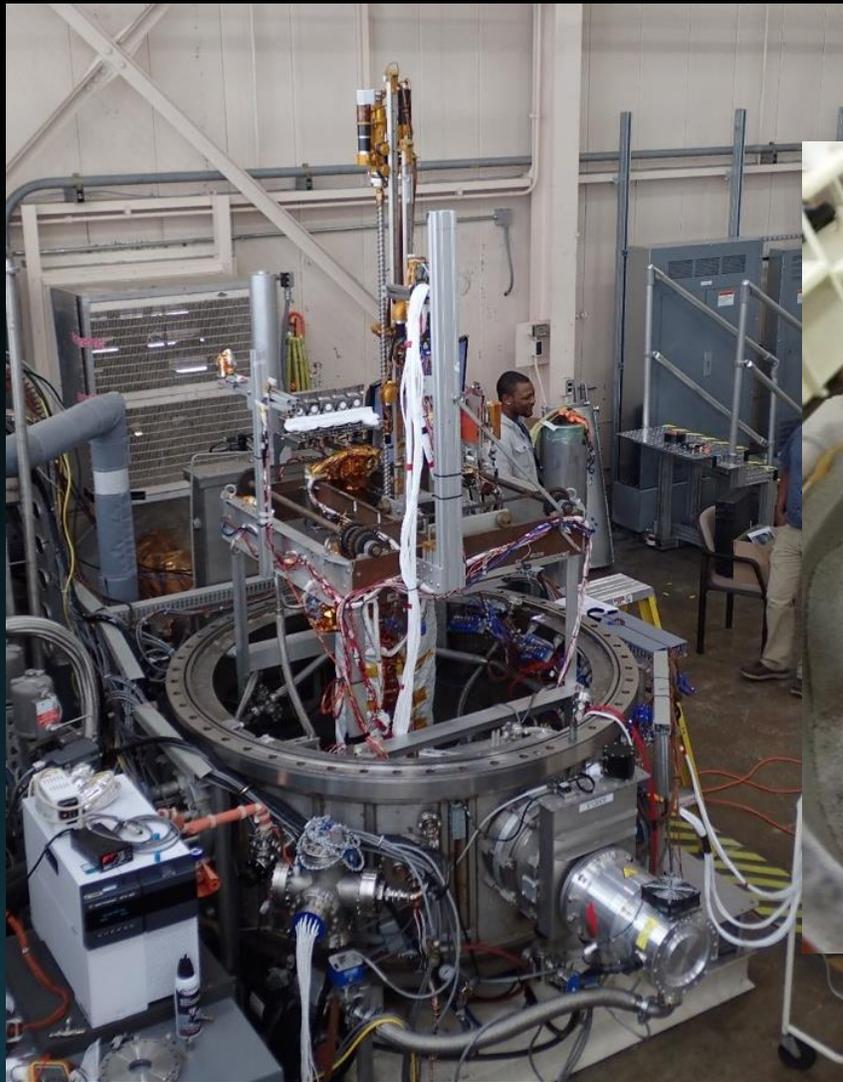
- **Analyzes volatiles excavated through rover traversing and drilling**
- Located under the rover studying water/volatiles abundance while roving & drilling

The Regolith and Ice Drill for Exploring New Terrain (TRIDENT) HBR¹

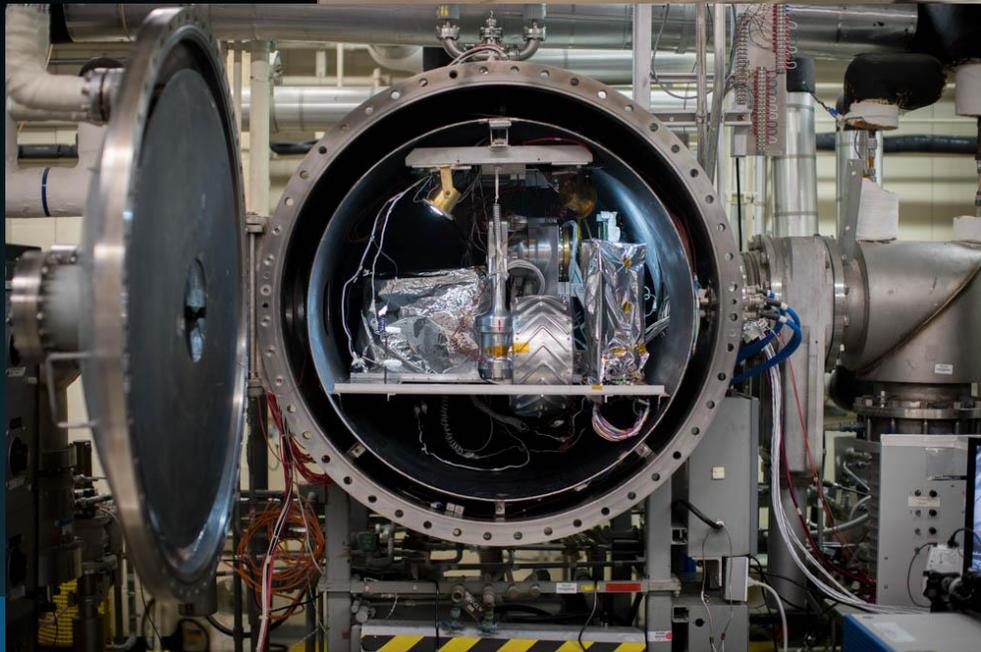
- **Excavates lunar regolith to a depth of 1-meter**, in 10cm increments
- Measures forces, displacements and temperatures for regolith bulk properties
- Located under the center of the rover to minimize volatiles solar sublimation

The image shows a close-up view of the lunar surface, characterized by numerous dark, circular craters of varying sizes. The surface is illuminated from the side, creating deep shadows and highlighting the rugged terrain. A solid, bright blue horizontal band runs across the center of the image, serving as a background for the title text.

VIPER Environmental Test



- Drill testing at NASA-GRC's VF-13 TVAC chamber
- Studying volatiles loss while drilling a meter deep into lunar soil
- Using engineered lunar-like soil conditions, doped with 5% water and chilled to -100C

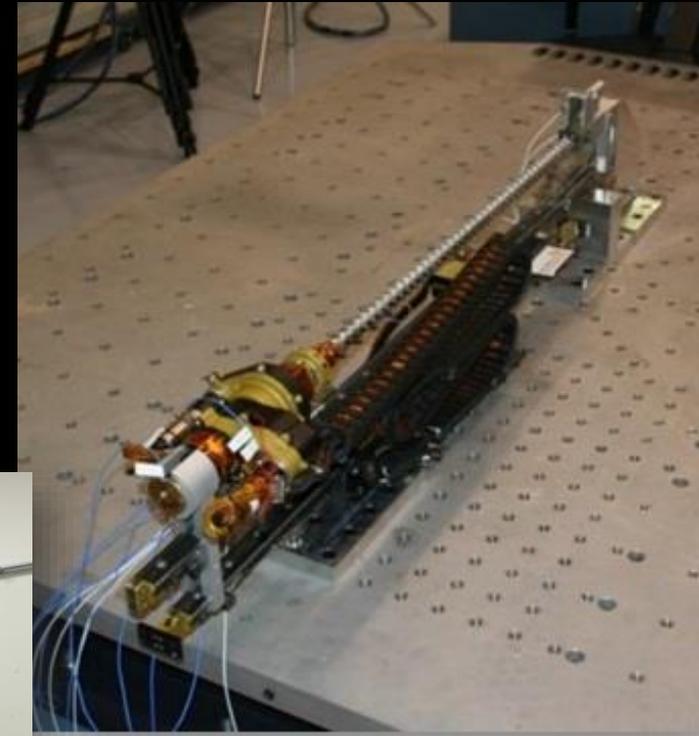


TVAC chamber testing of VIPER rover wheels & steering assemblies @ NASA-JSC



TVAC chamber Z-axis testing of VIPER rover

TVAC chamber Z-axis testing of VIPER rover



Drill undergoing Vibration testing



Studying impacts of the poor lighting and long shadows in polar regions

Field testing in the dark
- NASA-ARC Roverscape

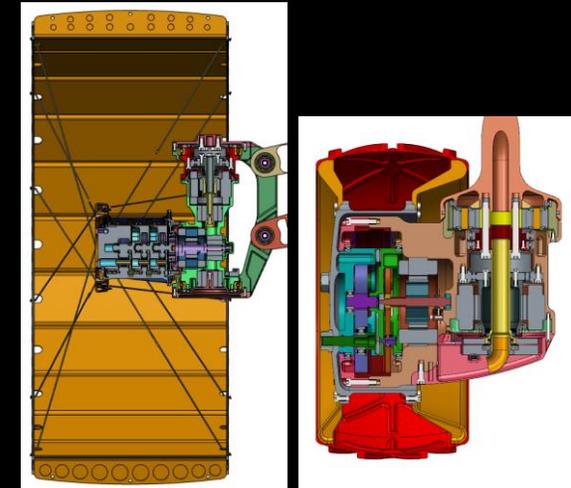


Crab-walking tests with the new 1A wheel. Studying traction performance while avoiding overlapping wheel tracks.

15° upslope with 45° cross slope



20° upslope

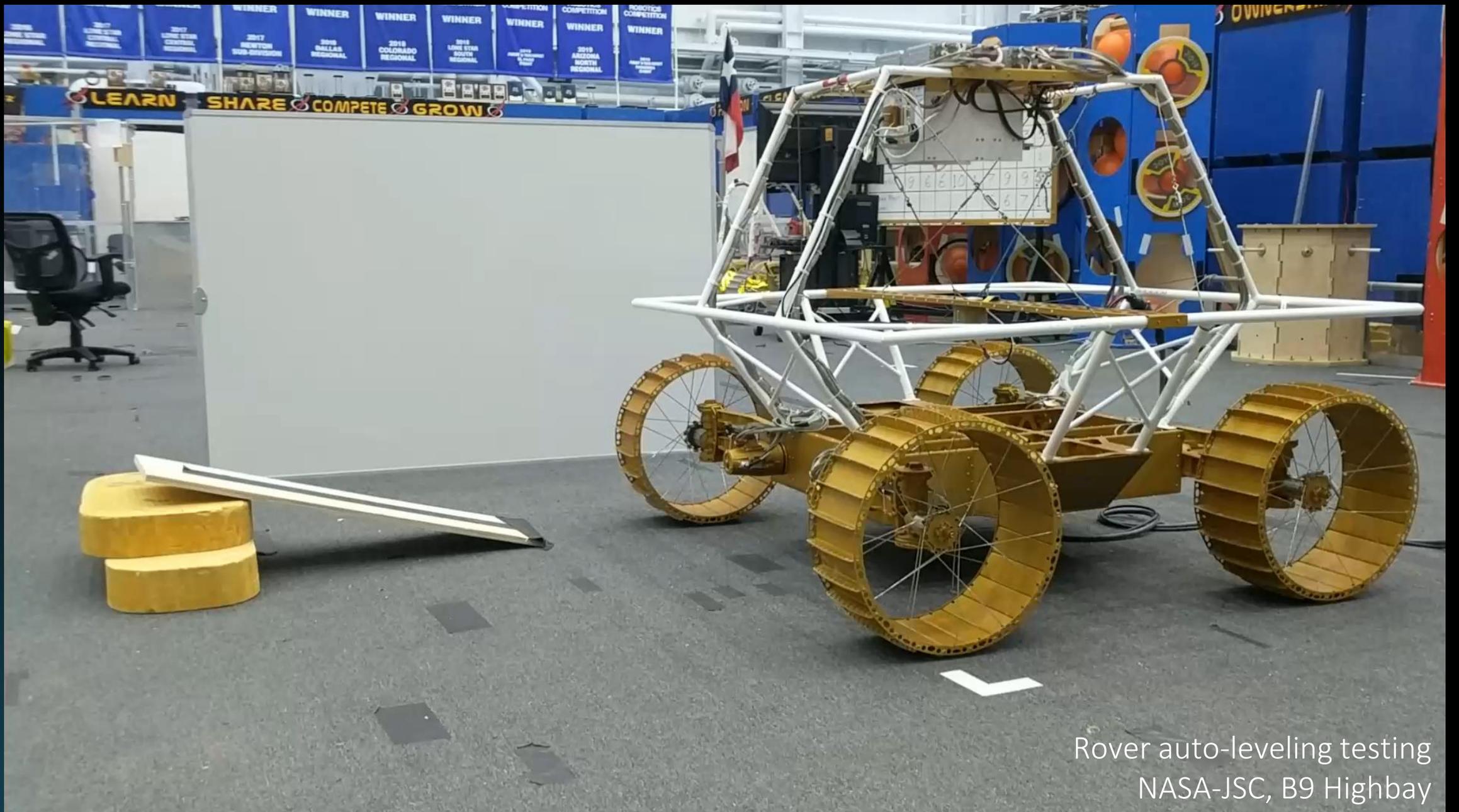


Rover undergoing soil environments testing

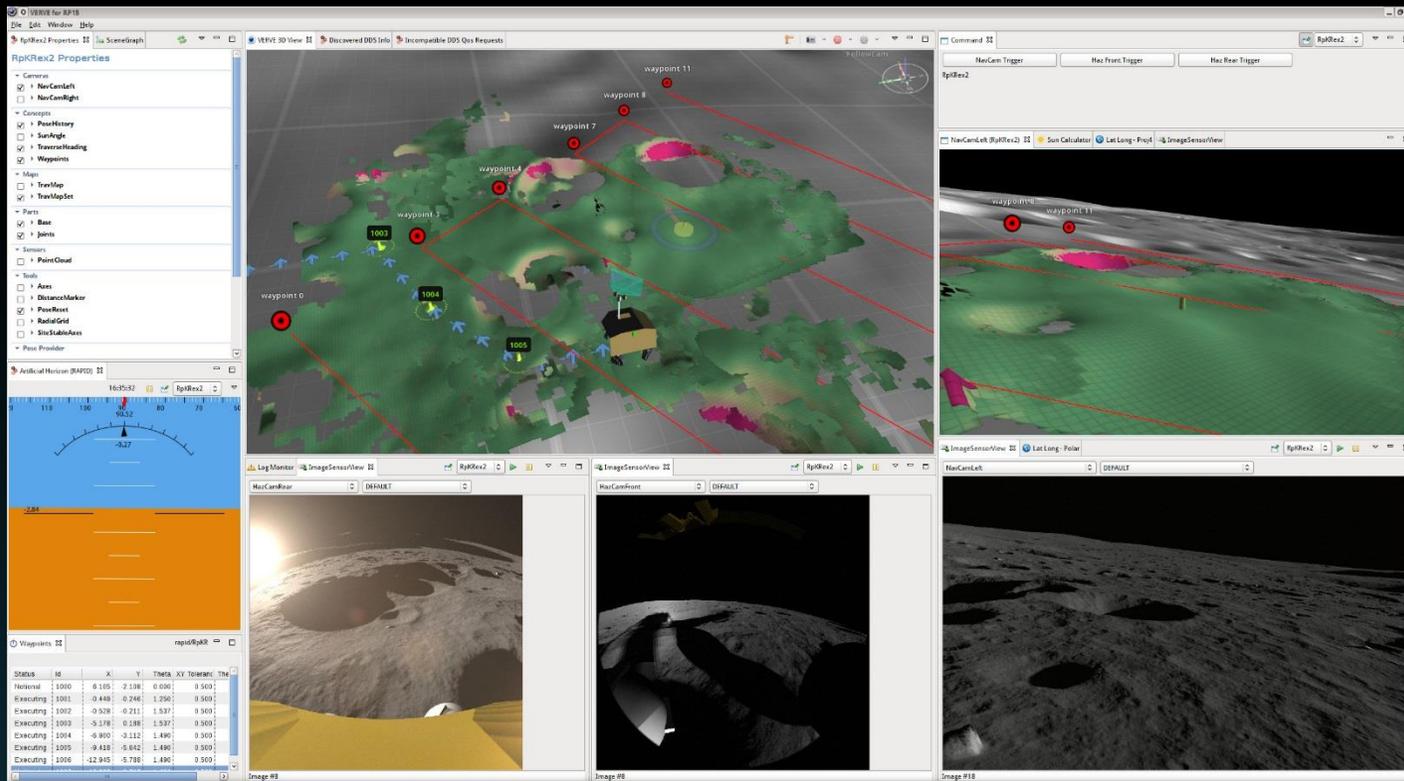
NASA-GRC Simulated Lunar Operations Lab (SLOPE)



Rover regolith simulant testing: NASA-KSC Regolith Lab



Rover auto-leveling testing
NASA-JSC, B9 Highbay



Powerful, fully-synthetic, lunar terrain sim based on Digital Elevation Maps (DEM)
 – Establishing driver decision-making times

Rover Driving Simulator Capability
 NASA-ARC Lunar Operations Lab



VIPER lunar-weight rover tested in lunar simulant soil bin
NASA-GRC, SLOPE laboratory



“Swimming” ...



x64

The image shows a close-up view of the lunar surface, characterized by numerous dark, circular craters of varying sizes. The surface is illuminated from the side, creating deep shadows and bright highlights. A solid teal horizontal band runs across the middle of the image, containing the word "Questions?" in white text.

Questions?

VIPER Neutron Spectrometer System (NSS)

NSS (NASA ARC, Lockheed Martin ATC)

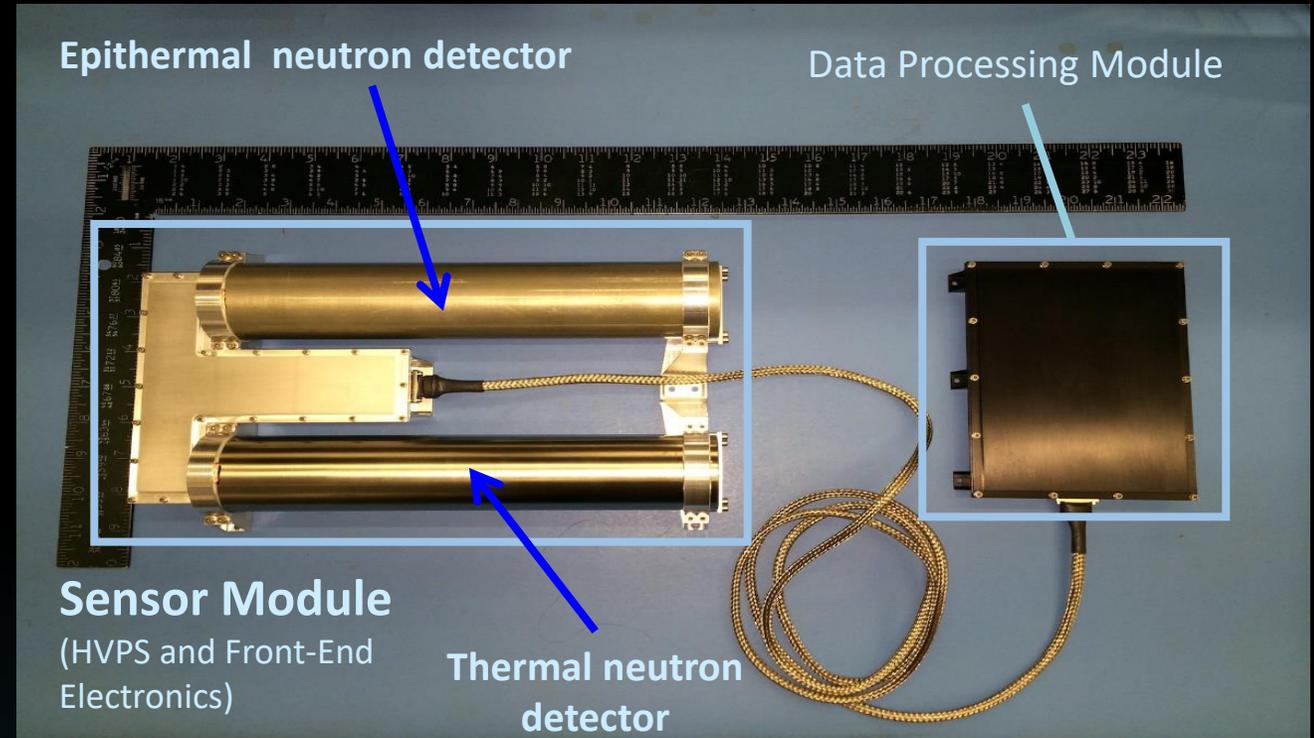
PI: Rick Elphic (NASA ARC)

Instrument Type: Two channel neutron spectrometer

Key Measurements: NSS assesses hydrogen and bulk composition in the top meter of regolith, measuring down to 0.5% (wt) WEH to 3-sigma while roving

Operation: On continuously while roving

Instrument Name	NSS
Mass [kg], CBE	1.9*
Dimensions [cm]	Sensor Module: 21.3 x 32.1 x 6.8 Data Processing Module: 13.9 x 18.0 x 3.0
Power [W]	1.6
Sensitivity	WEH to ≥ 0.5 wt% water-equivalent at 10 cm/s
Accuracy	5 – 10% absolute



VIPER Near InfraRed Volatiles Spectrometer System (NIRVSS)

NIRVSS (ARC, Brimrose Corporation)

PI: Anthony Colaprete (NASA ARC)

Instrument Type: NIR Point Spectrometer, 4Mpxl Panchromatic Imager with 7 LEDs, four channel thermal radiometer

Key Measurements: Volatiles including H₂O, OH, and CO₂ and, mineralogy, surface morphology and temperatures

Operation: On continuously while roving and during drill operations

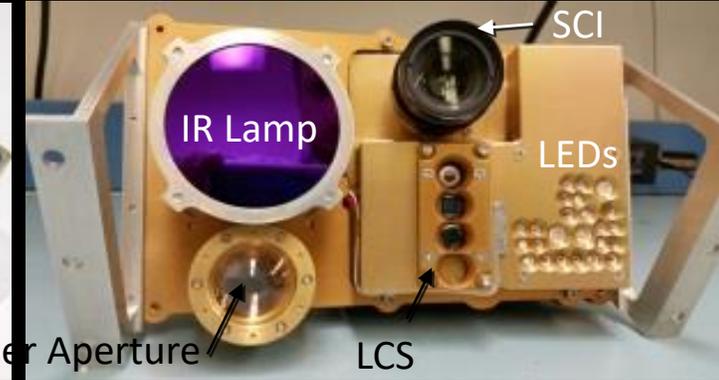
Primary Measurements:

- *AOTF NIR Point Spectrometer:* 1300-4000nm
- *Spectrometer Context Imager (SCI):* 4Mpxl imager with seven LEDs between 340-940nm
- *Longwave Calibration Sensor (LCS):* IR flux and surface temperature down to <100K to ± 5K
- *Lamp:* Dual filament tungsten lamp provides even, calibrated light source when in shadow

Spectrometer



Bracket Assembly



Instrument Name	NIRVSS
Mass [kg]	3.57 kg (not including Fiber)
Dimensions [cm]	Spectrometer Module: 18x18x8.5 Observation Bracket: 20.4x13x15.1
Power [W], Avg	Spectrometer = 12 Bracket Assembly = 5.26 Lamp = 12.3
Sensitivity	Range: 1.2 to 4.0 mm SNR>100 at 2 and 3 mm Water Ice to <0.25%
Accuracy	Radiance to <25%

VIPER Mass Spectrometer Observing Lunar Operations (MSolo)

MSolo (KSC, INFICON)

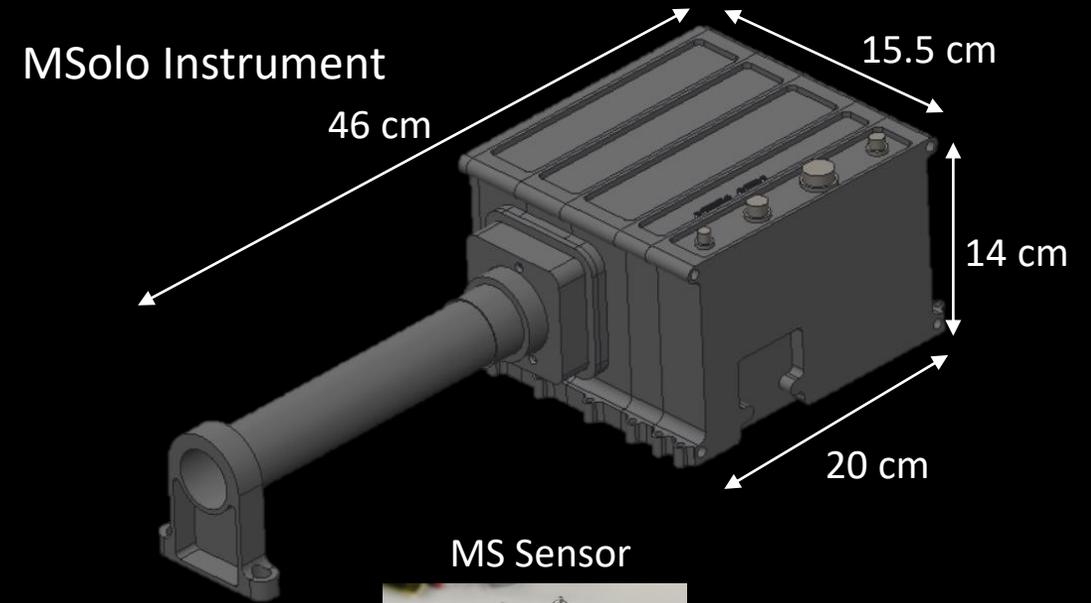
PI: Janine Captain (NASA KSC)

Instrument Type: Quadrupole mass spectrometer

Key Measurements: Identify low-molecular weight volatiles between 2-100 amu, unit mass resolution to measure isotopes including D/H and O^{18}/O^{16}

Operation: Views below rover and at drill cuttings, volatile analysis while roving and during drill activities

Instrument Name	MSolo
Mass, CBE	6 kg
Dimensions	15.5 x 20 x 46 cm
Power	Average 35 W while scanning
Detectors	Faraday Cup (MDPP* $1.5e-12$ Torr) Electron Multiplier (MDPP* $2e-15$ Torr)



*MDPP – minimum detectable partial pressure @ m/z 28 with open ion source

VIPER The Regolith and Ice Drill for Exploring New Terrain (TRIDENT) Drill

TRIDENT (Honeybee Robotics)

PI: Kris Zacny (Honeybee)

Instrument Type: 1-meter hammer drill

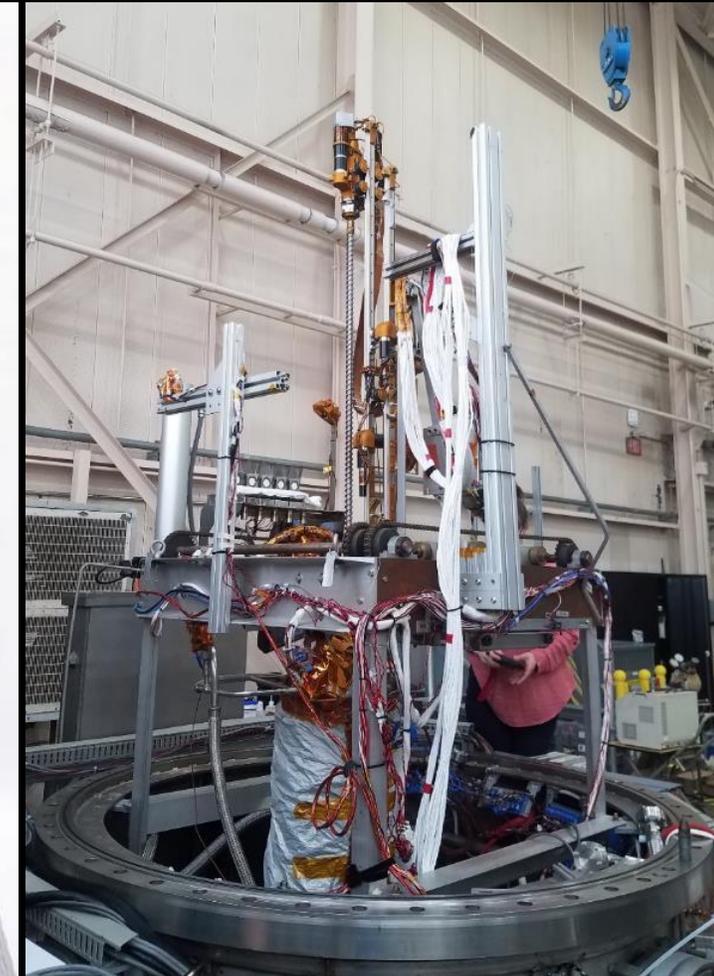
Key Measurements: Excavation of subsurface material to 100 cm; Subsurface temperature vs depth; Strength of regolith vs depth (info on ice-cemented ground vs. ice-soil mixture).

Operation: Performs subsurface assays down to 100 cm in <1 hr, depositing cuttings at surface for inspection

Instrument Name	TRIDENT
Mass [kg], CBE	18 (includes launch locks). Can be reduced for lander deployment.
Dimensions (stowed) [cm]	27 x 22 x 177 (for 1-m depth). Can be reduced for lander deployment.
Power [W]	Idle: < 5 Augering: ~20 nominal, 175 max Percussion: 0 nominal, 150 max
Telemetry (while operating)	~3.4 kbits/s



TRL6 Drill



Lunar cryo-chamber tests at GRC